

Advanced Higher Physics

Past Paper Questions

1.6 Stellar Physics

THE VERY LARGE TELESCOPE	<input checked="" type="checkbox"/>
THE EXTREMELY LARGE TELESCOPE	<input checked="" type="checkbox"/>
THE OVERWHELMINGLY LARGE TELESCOPE	<input checked="" type="checkbox"/> (CANCELED)
THE OPPRESSIVELY COLOSSAL TELESCOPE	<input type="checkbox"/>
THE MIND-NUMBINGLY VAST TELESCOPE	<input type="checkbox"/>
THE DESPAIR TELESCOPE	<input type="checkbox"/>
THE CATAclysmic TELESCOPE	<input type="checkbox"/>
THE TELESCOPE OF DEVASTATION	<input type="checkbox"/>
THE NIGHTMARE SCOPE	<input type="checkbox"/>
THE INFINITE TELESCOPE	<input type="checkbox"/>
THE FINAL TELESCOPE	<input type="checkbox"/>

6. A Hertzsprung-Russell (H-R) diagram is shown in Figure 6A.

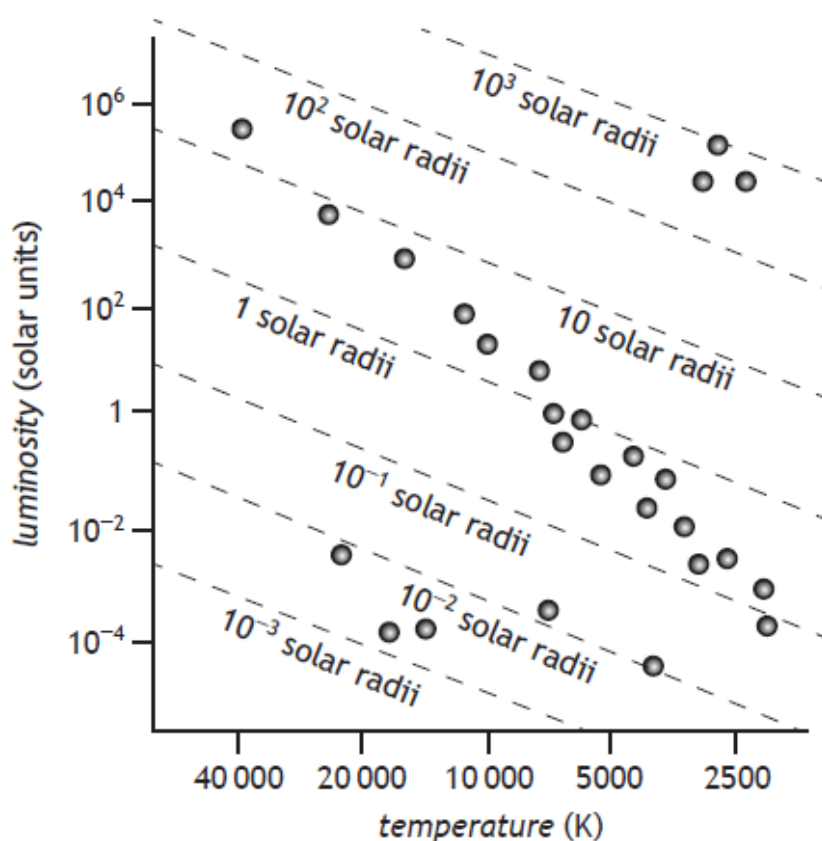
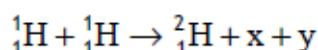


Figure 6A

- (a) All stars on the main sequence release energy by converting hydrogen to helium. This process is known as the proton-proton (p-p) chain. One stage in the p-p chain is shown.



Name particles x and y .

2

- (b) The luminosity of the Sun is $3.9 \times 10^{26} \text{ W}$. The star Procyon B has a luminosity of 4.9×10^{-4} solar units and a radius of 1.2×10^{-2} solar radii.

(i) On the H-R diagram, circle the star at the position of Procyon B.
(An additional diagram, if required, can be found on Page 42.)

1

(ii) What type of star is Procyon B?

1

(iii) The apparent brightness of Procyon B when viewed from Earth is $1.3 \times 10^{-12} \text{ W m}^{-2}$.

Calculate the distance of Procyon B from Earth.

4

(c) The expression

$$\frac{L}{L_0} = 1.5 \left(\frac{M}{M_0} \right)^{3.5}$$

can be used to approximate the relationship between a star's mass M and its luminosity L .

L_0 is the luminosity of the Sun (1 solar unit) and M_0 is the mass of the Sun.

This expression is valid for stars of mass between $2M_0$ and $20M_0$.

Spica is a star which has mass $10.3M_0$.

Determine the approximate luminosity of Spica in solar units.

2

2016

4. Epsilon Eridani is a star 9.94×10^{16} m from Earth. It has a diameter of 1.02×10^9 m. The apparent brightness of Epsilon Eridani is measured on Earth to be $1.05 \times 10^{-9} \text{ W m}^{-2}$.

(a) Calculate the luminosity of Epsilon Eridani.

3

(b) Calculate the surface temperature of Epsilon Eridani.

3

2015

4. The constellation Orion, shown in Figure 4A, is a common sight in the winter sky above Scotland.

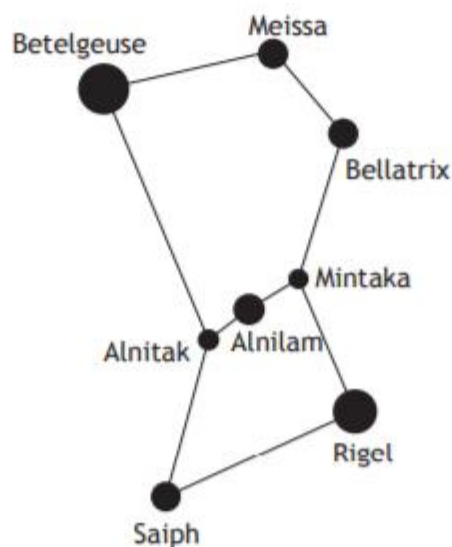
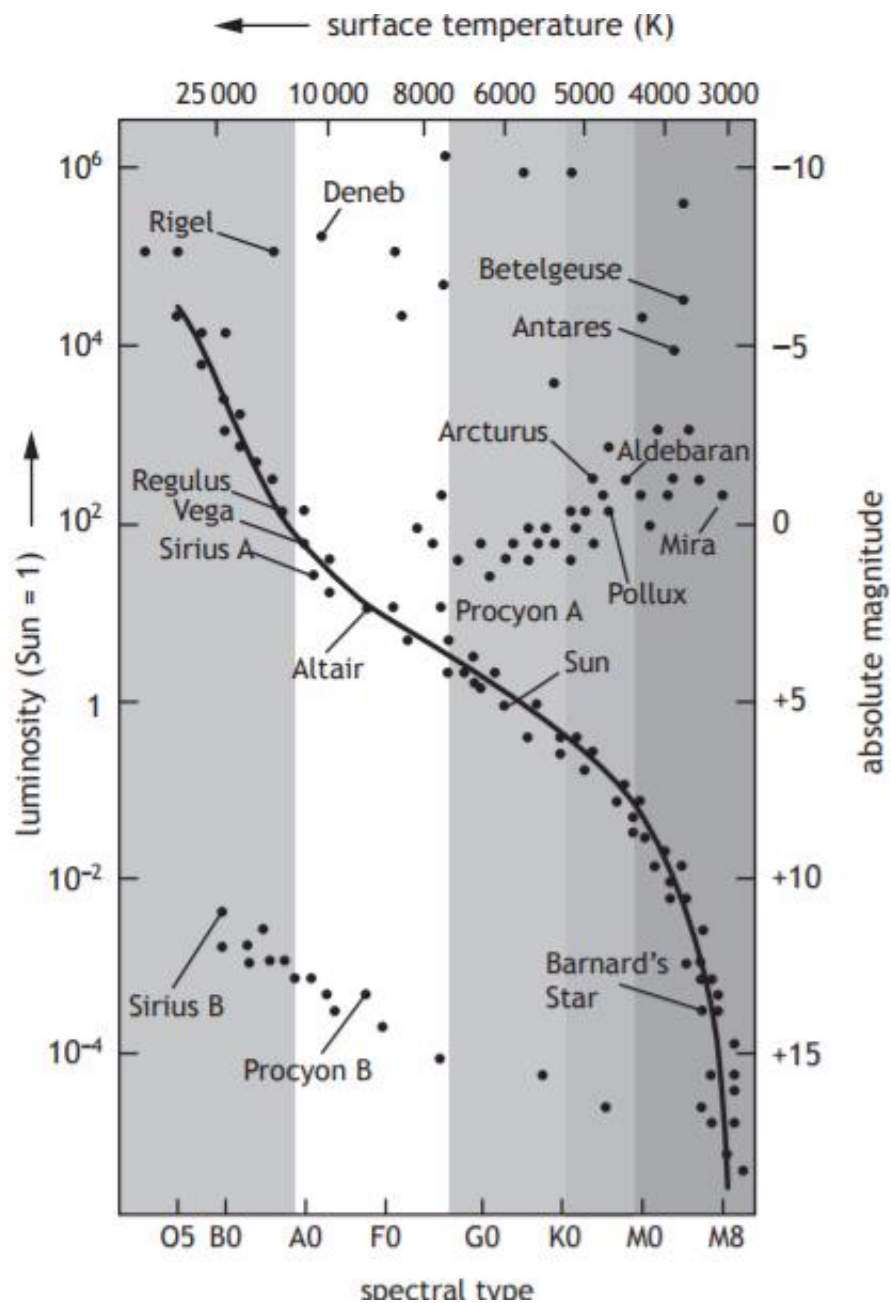


Figure 4A

Past

Two of the stars in this constellation are known as Betelgeuse and Rigel. Their positions are shown on the Hertzsprung-Russell (H-R) diagram in Figure 4B.

2



(a) Using the H-R diagram, predict the colour of Betelgeuse.

1

(b) The table shows some of the physical properties of Rigel.

<i>Property of Rigel</i>	
Surface temperature	$(1.20 \pm 0.05) \times 10^4 \text{ K}$
Radius	$(5.49 \pm 0.50) \times 10^{10} \text{ m}$
Mass	18 ± 1 solar masses
Distance to Earth	773 ± 150 light years

(i) (A) Calculate the luminosity of Rigel.

3

- (B) State the assumption made in your calculation. 1
- (ii) Calculate the absolute uncertainty in the value of the luminosity of Rigel. 4
- (c) Calculate the apparent brightness of Rigel as observed from the Earth. 4
- (d) Betelgeuse is not on the Main Sequence region of the H-R diagram. Describe the changes that have taken place in Betelgeuse since leaving the Main Sequence. 2

Revised Advanced Higher 2015

3. The luminosity of the Sun is 3.9×10^{26} W and the mean radius of the Earth's orbit around the Sun is 1 astronomical unit (AU).
- (a) Calculate the Sun's apparent brightness at the surface of the Earth. 2
- (b) The distance d to a star can be calculated using the relationship:

$$10^{0.2(m-M)} = \frac{d}{10}$$

This gives a distance in parsecs. 1 parsec is equivalent to 3.26 light years.

The apparent magnitude (m) of a celestial body is a measure of its brightness as viewed from Earth. The absolute magnitude (M) of a celestial body is a measure of its intrinsic brightness.

The following data was obtained for a star.

$$\begin{aligned} \text{Apparent magnitude} &= 5.62 \\ \text{Absolute magnitude} &= -4.38 \end{aligned}$$

- Calculate the distance in light years to this star from Earth. 3
- (5)**

4. The lyrics of the song Woodstock contain the lines "We are stardust; we are golden. We are billion year old carbon".
- Use your knowledge of Physics to comment on these lyrics. (3)

5. A typical Hertzsprung-Russell (H-R) diagram is shown in Figure 5A.

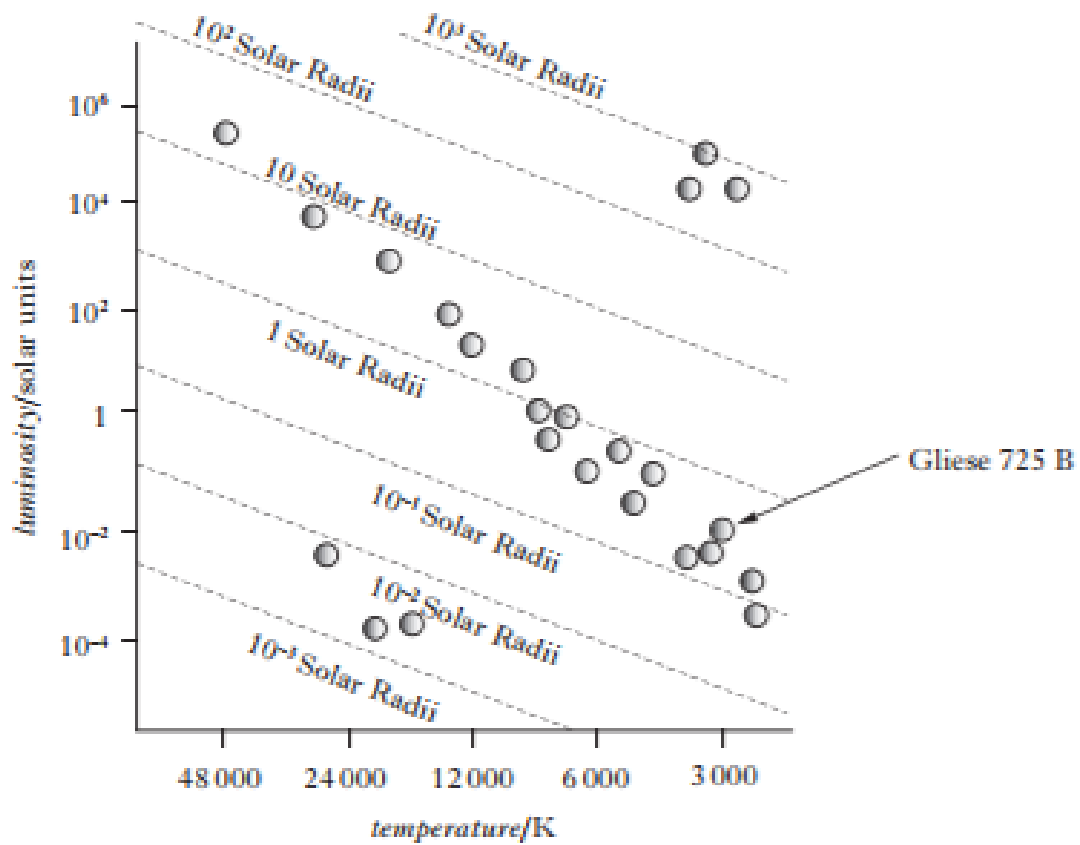


Figure 5A

(a) The luminosity of the Sun is 3.9×10^{26} W.

Using information from Figure 5A:

- (i) determine the luminosity in watts of Gliese 725 B; 1
- (ii) show that the radius of Gliese 725 B is 3×10^8 m; 2
- (iii) explain why it would be inappropriate to give the answer for part (ii) to more than one significant figure. 1

- (b) Figure 5B shows how the radiation intensity varies with frequency for a black body radiator.

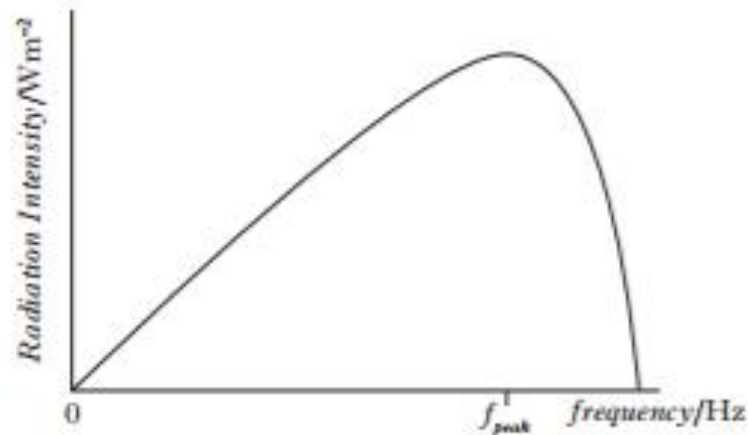


Figure 5B

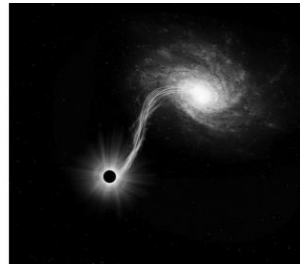
This spectrum has a peak intensity at a frequency of f_{peak} . f_{peak} can be estimated using the relationship

$$f_{peak} = \frac{2.8k_b T}{h}$$

where $k_b = 1.38 \times 10^{-23} \text{ J K}^{-1}$ (Boltzmann constant) and the other symbols have their usual meanings.

- (i) Estimate f_{peak} for Gliese 725 B. 2
- (ii) The cosmic microwave background radiation (CMBR) has a spectrum which peaks at a wavelength of 1.9 mm. Calculate the temperature of the CMBR. 3
- (c) Some astronomers have suggested that primordial black holes of mass 1.0×10^{-10} solar masses could make up the dark matter in our galaxy. Determine the Schwarzschild radius of such a black hole. 3
- (12)**

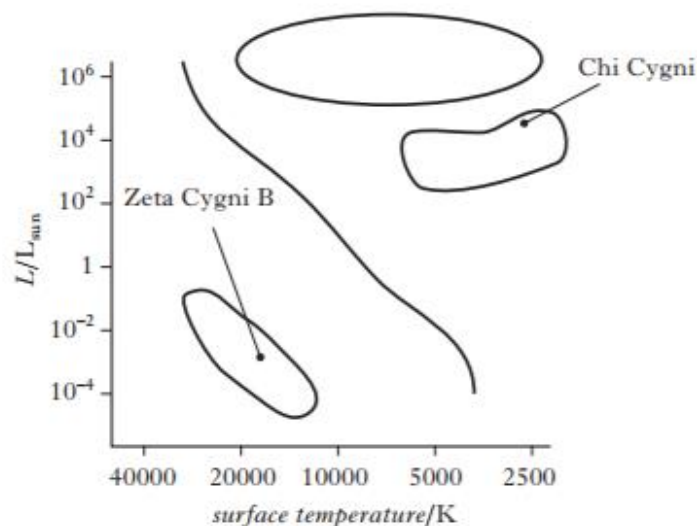
4. Cygnus X-1 is an X-ray source in the constellation Cygnus that astrophysicists believe contains a black hole. An artist's impression is shown in Figure 4A.



The mass of the black hole has been determined to be 14.8 Solar masses.

- (a) (i) State what is meant by the Schwarzschild radius of a black hole. 1
 (ii) Calculate the Schwarzschild radius of the black hole in Cygnus X-1. 3

- (b) The Hertzsprung-Russell (H-R) diagram shown in Figure 4B shows the relationship between the luminosity and surface temperature of stars.



Zeta Cygni B and Chi Cygni are two stars in the constellation Cygnus. They are shown on the H-R diagram. Chi Cygni is more luminous than Zeta Cygni B.

Describe **two** other differences between these stars. 2

- (c) Another star, Aldebaran B, is a distance of 6.16×10^{17} m from the Earth.

The luminosity of Aldebaran B is 2.32×10^{25} W and its temperature is determined to be 3.4×10^3 K.

- (i) Calculate the radius of Aldebaran B. 2
 (ii) Calculate the apparent brightness of Aldebaran B as observed from Earth. 2

5. Hertzsprung-Russell (H-R) diagrams are widely used by physicists and astronomers to categorise stars. Figure 5A shows a simplified H-R diagram.

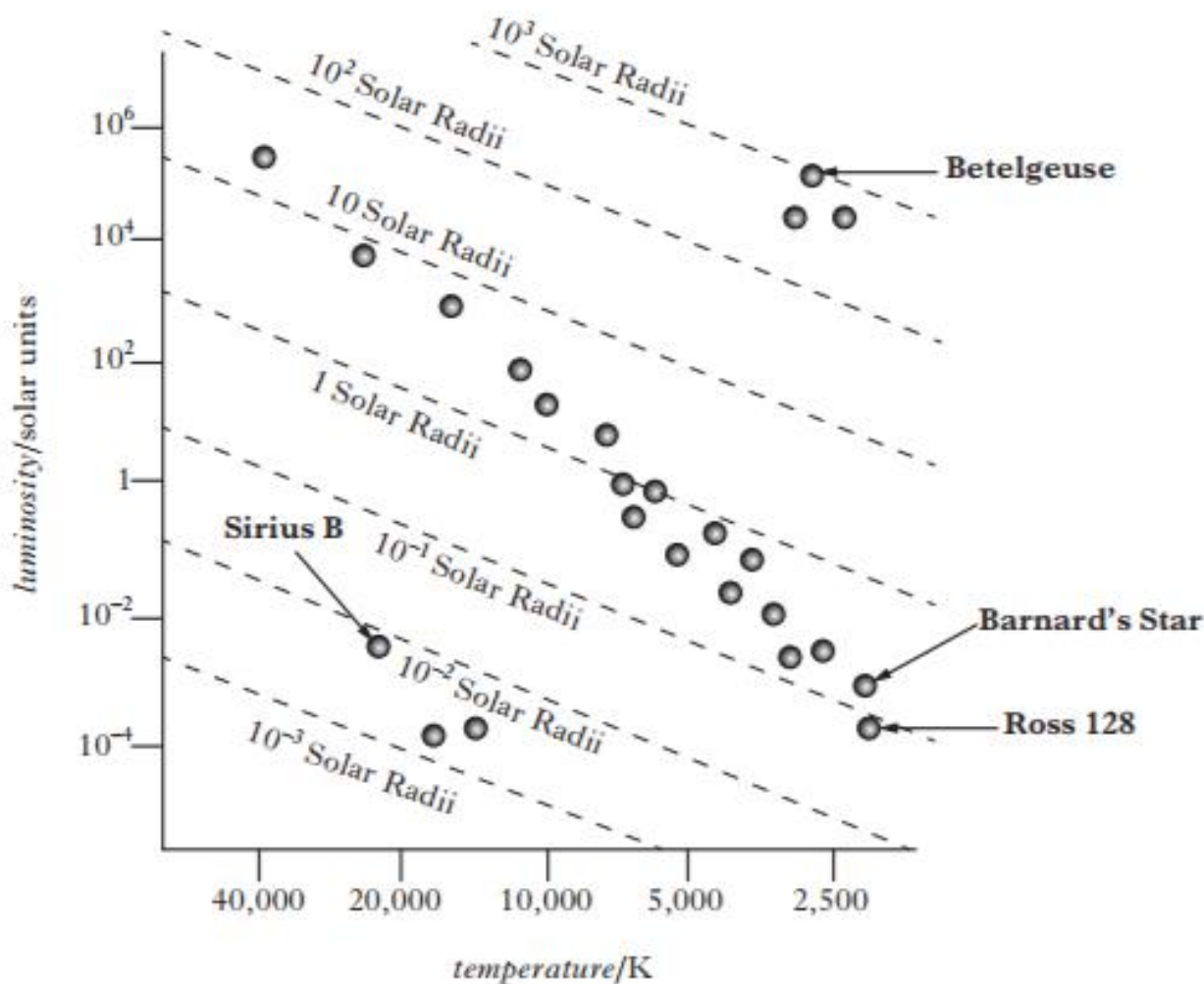


Figure 5A

- (a) What class of star is Sirius B? 1
- (b) Estimate the radius in metres of Betelgeuse. 2
- (c) Ross 128 and Barnard's Star have a similar temperature but Barnard's Star has a slightly greater luminosity. What other information does this tell you about the two stars? 1

- (d) During the life cycle of the Sun its position in the H-R diagram is expected to change as shown by the arrowed line in Figure 5B.

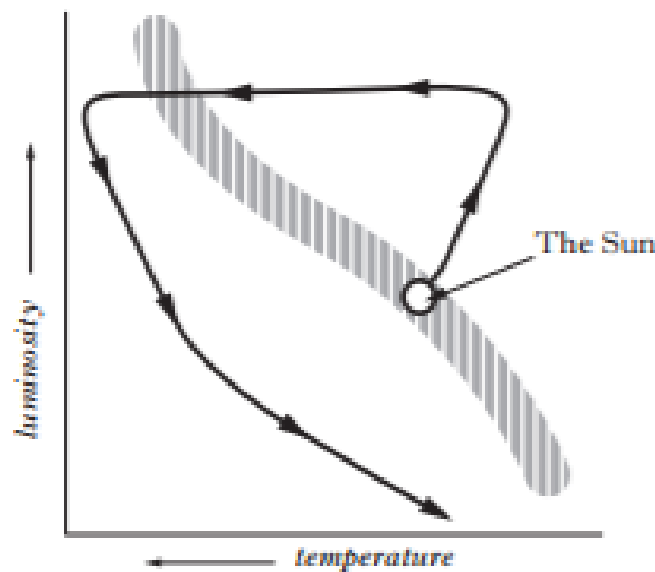


Figure 5B

Describe the changes that occur to the Sun during its expected life cycle. 3

- (e) Hydrogen fusion in a star is a result of a proton-proton chain. The process eventually results in the production of a helium-4 nucleus.
- Show that the percentage loss of mass from four protons to one helium-4 nucleus is 0.7%. 2
 - The luminosity of the Sun is 3.8×10^{26} W. Using Einstein's energy equation, show that the mass of hydrogen lost per second in the Sun is 4.2×10^9 kg. 1
 - Estimate the lifetime of the Sun in seconds. Assume the mass of hydrogen in the Sun to be the same as the mass of the Sun. 1
- (f) The "no greenhouse" temperature of a planet is the average surface temperature of a planet in the absence of any greenhouse effect. The "no greenhouse" temperature of a planet in kelvin is given by

$$T = 280 \left(\frac{(1 - \text{reflectivity})}{d^2} \right)^{\frac{1}{4}}$$

where d is the distance from the Sun in astronomical units (AU).

The reflectivity is a measure of the percentage of energy reflected from the surface, 1 represents 100% reflectivity and 0 represents no reflectivity.

Mercury has a reflectivity of 0.12 and is 5.8×10^{10} m from the Sun.

Calculate its "no greenhouse" temperature. 2

(13)

